



## Radiometric System Model for RBI

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### Radiometric System Model Objectives



#### Develop a tool to enhance the interpretation of Instrument performance

Model the end-to-end science signal chain: Photons in to bits/counts out.

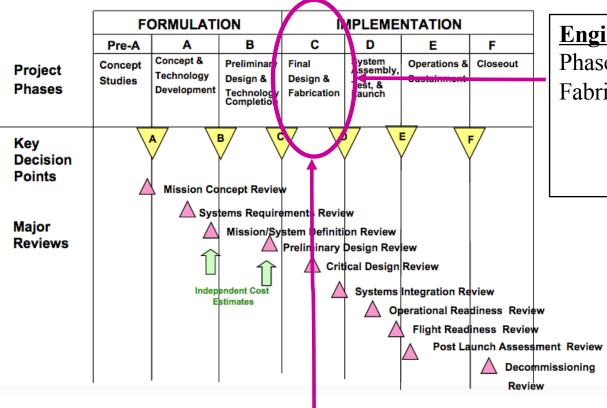


- Simulate the science data stream output when viewing calibration targets, earth scenes or any user-defined radiance.
- Support and validate engineering design and fabrication phase
- Quantify the effects of various anomalous sources of energy: stray light
- Perform analyses as required to evaluate and quantify the impact to science data due to other uncertainties.



### **Current RBI Project Phase**





#### **Engineering-Led Effort**

Phase C: Final Design and Fabrication

 Demonstrate that the detailed system design meets requirements

#### **Science-Led Effort**

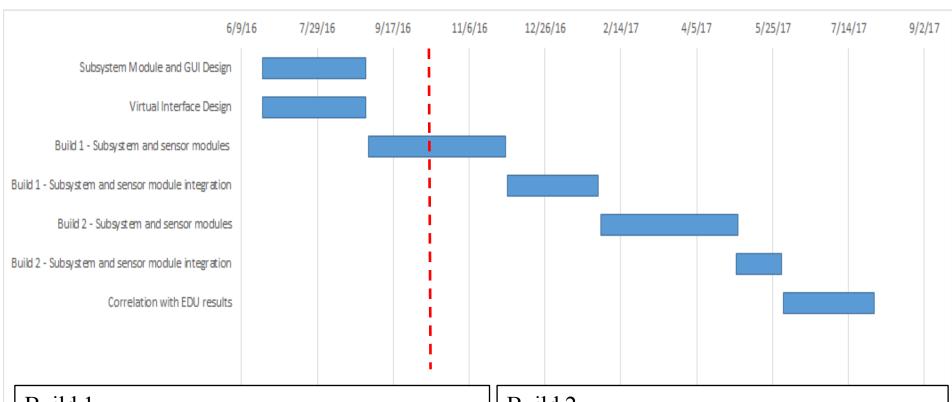
Develop the end-to-end model of the science signal chain: Photons in to bits/counts out.

- To be correlated to the Engineering Development Unit at the end of Phase C
- To be correlated to the Flight Unit at the end of Phase
- Support Mission Operations and Data Analysis in Phase E



### **Model Development Schedule**





#### Build 1

Individual modules complete Monochromatic sources Single telescope (Total) Single sided GUI

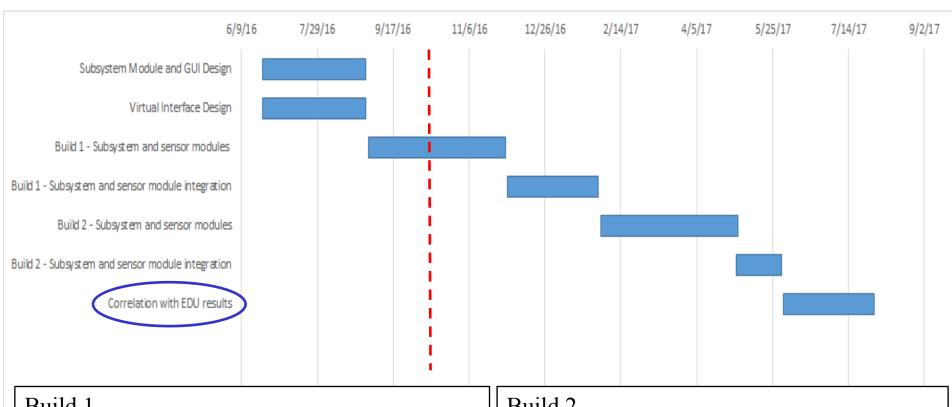
#### Build 2

Broadband sources All Three telescopes Double sided



### **Model Development Schedule**





#### Build 1

Individual modules complete Monochromatic sources Single telescope (Total) Single sided GUI

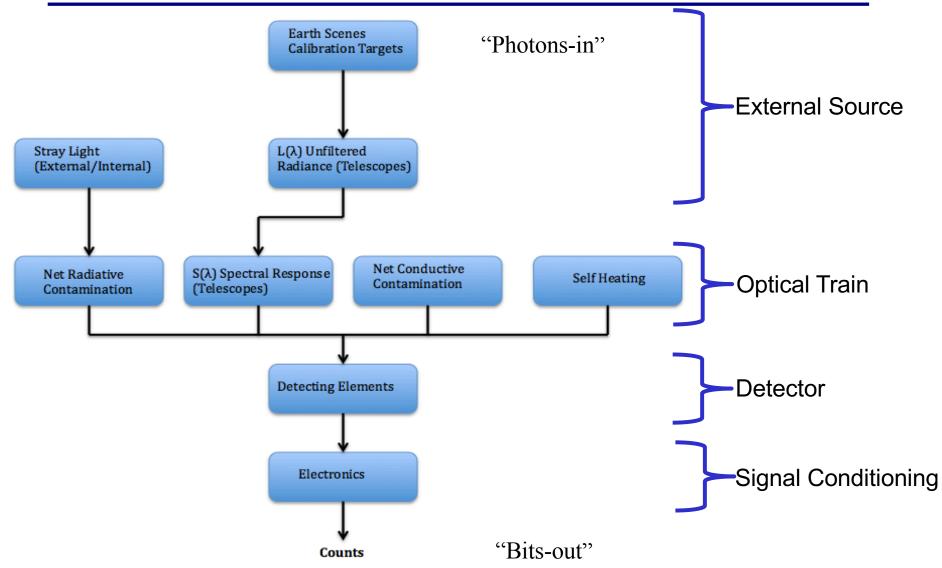
#### Build 2

Broadband sources All Three telescopes Double sided



### **Typical Instrument Analytical Model**







## **Numerical Modeling Tools**



- Monte-Carlo Ray-Trace Model
  - Computes the distribution of radiation within the instrument.
  - Spectral characterization of the optical and radiative performance of the entire instrument.
  - Provides the necessary "Boundary" conditions for the thermal models.
- Finite-Element Thermal Diffusion Model
  - Three-Dimensional characterization of the transient thermal diffusion in instrument components
- Finite-Difference Electro-thermal Model
  - Three-Dimensional characterization of the transient thermal diffusion in the detectors
  - Two-Dimensional characterization of the transient electrical diffusion in the thermocouples.
- Electrical Circuit Model
  - Computation of the electronic Response of the electrical feedback control system.

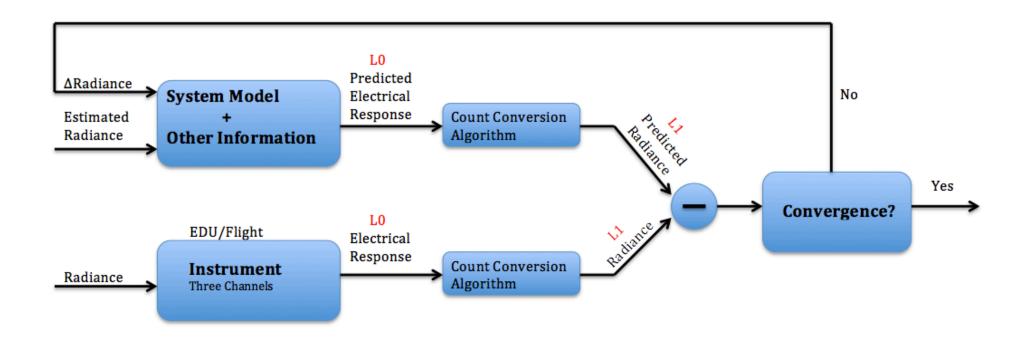
Previous Earth Radiation Budget (ERB) programs, such as CERES, have used these modeling tools for End-to-End characterization of the instrument



### **Correlation of Model to Hardware**



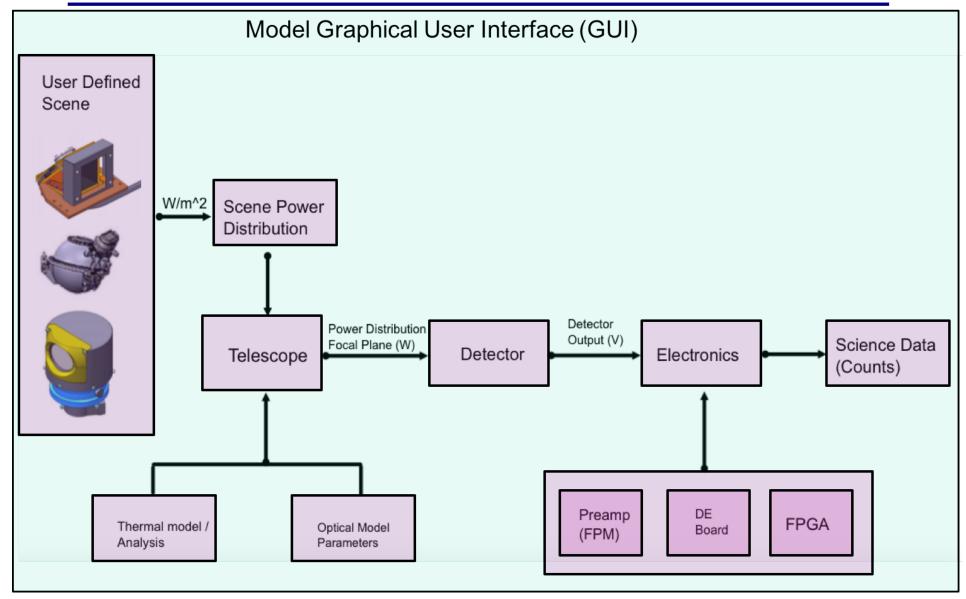
- During System Level TVAC testing we will simulate the test execution with the model to complete an end-to-end correlation.
- If the Model and Hardware do not converge, we will perturb model parameters within their allowed tolerances to bring the model and hardware into agreement.





### **RBI Instrument End-to-End Model**

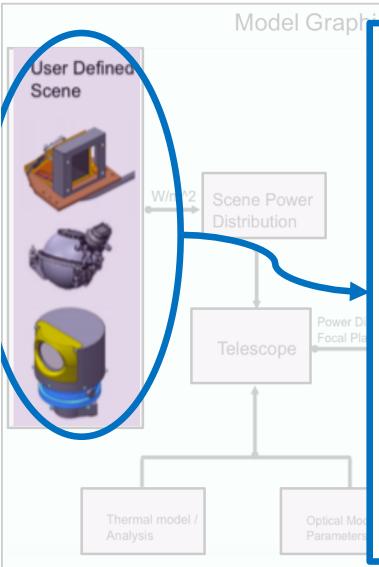






### **Sources**





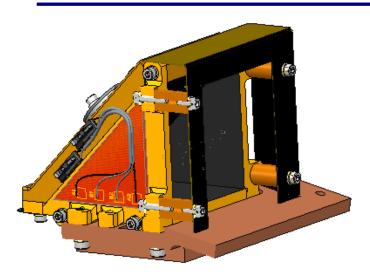
- On-board Calibration sources are currently being modeled with MATLAB and/or Zemax Optics Studio
  - Infrared Calibration Target (ICT) Using MCRT techniques to compute the
    distribution of radiation within the ICT
    and as distributed on the telescope
    aperture.
  - Visible Calibration Target (VCT) -Similar techniques as the ICT
  - Solar Calibration Target (SCT) –
     Similar techniques as the ICT
  - Parameters such as optical prescription, viewing geometry, and paint specs are also being modeled.
  - Thermal analysis are being conducted in parallel.

to

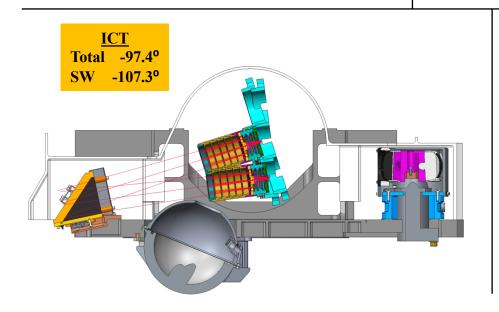


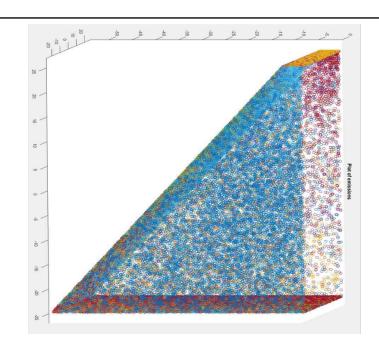
## **Infrared Calibration Target Module**





- ICT positioned to be viewed by both Total and Longwave Detectors
- ICT spatial and spectral output distribution imaged on the Focal Plane
- Thermal gradients within the ICT can produce ambiguous radiance
- Degradation of Z-302 will reduce effective emissivity over lifetime

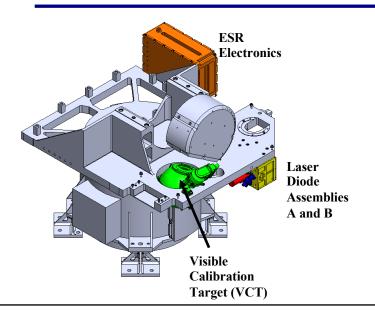




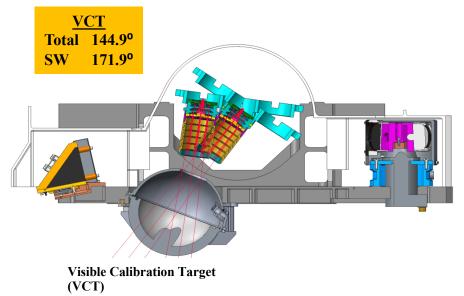


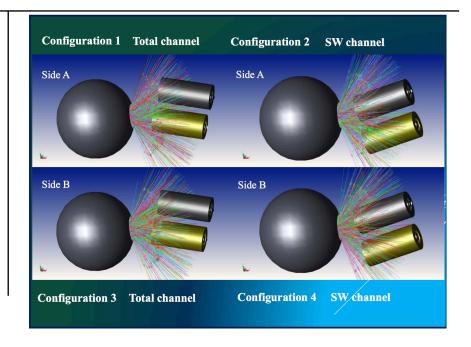
### **Visible Calibration Target Module**





- VCT positioned to be viewed by both Total and Shortwave Detectors
- VCT spatial and spectral output distribution imaged on the Focal Plane
- Thermal gradients within the VCT can produce IR background signal on the Total Channel

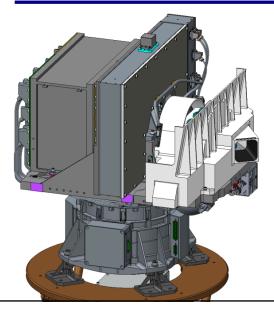




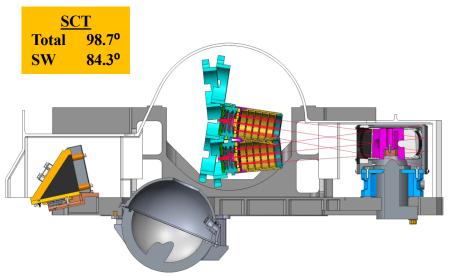


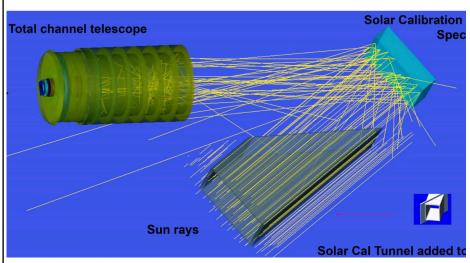
## **Solar Calibration Target Module**





- SCT positioned to be viewed by both Total and Shortwave Detectors
- SCT spatial and spectral reflected radiance distribution imaged on the Focal Plane
- Thermal gradients across the SCT can produce IR background signal on the Total Channel

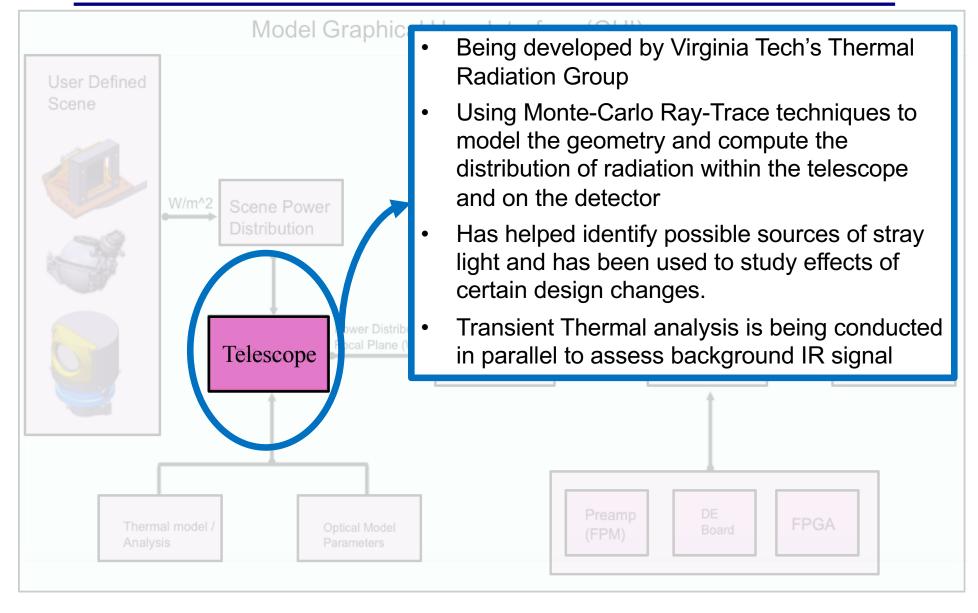






### **Optical Module**



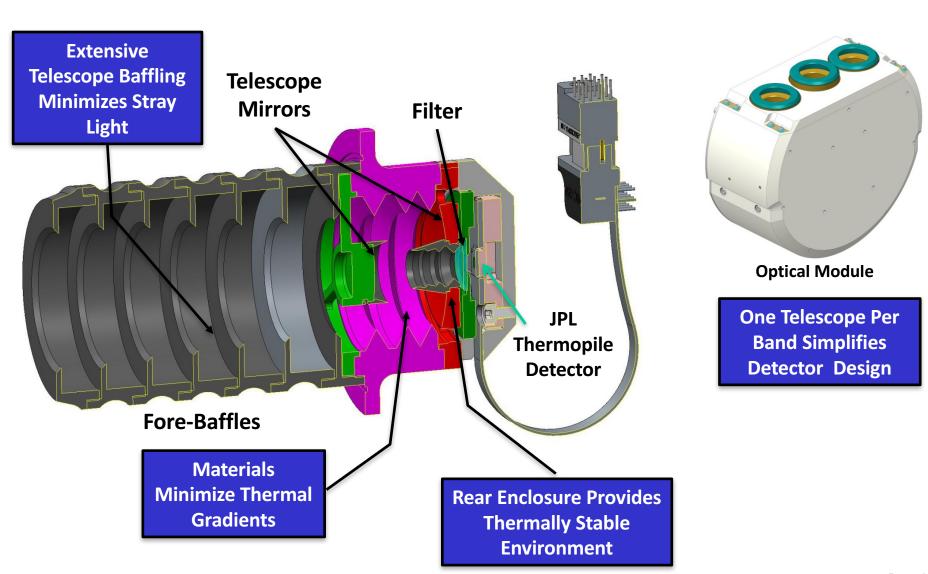




## **Optical Module**



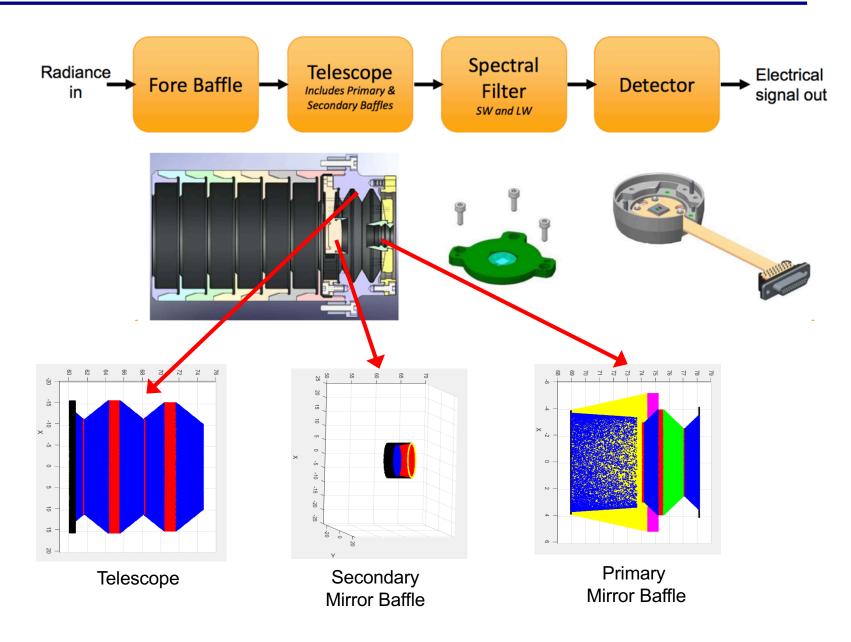
Radiation Budget Instrument





## **Optical Module**

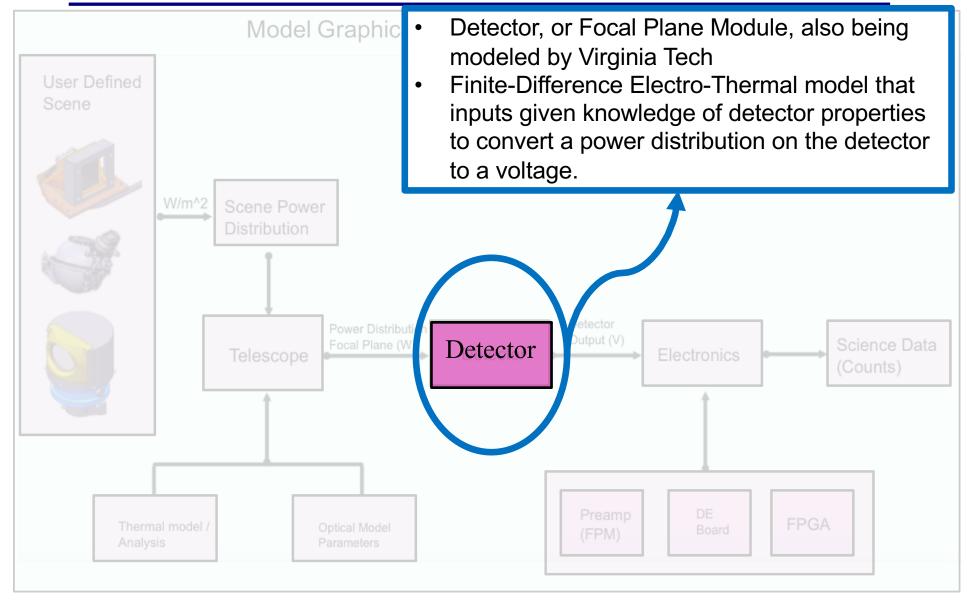






### **Focal Plane Model**

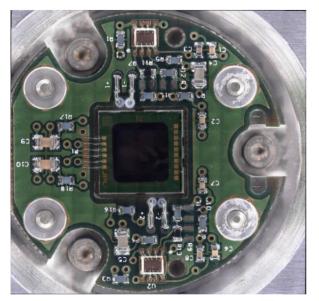


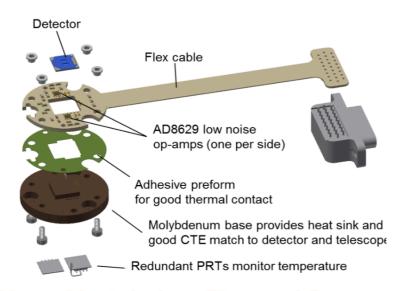


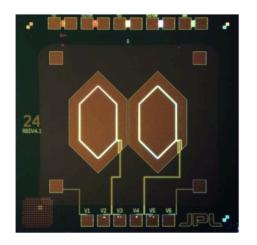


### **Focal Plane Module**









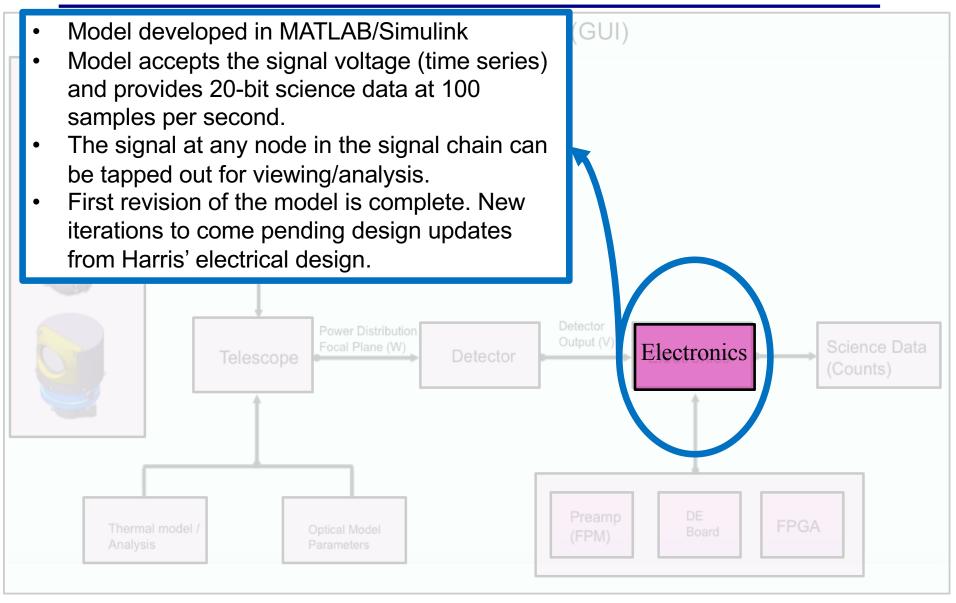
Focal Plane Module is a Thermal Detector that compares heat sink temperature with a membrane coupled to the radiation from the telescope

Passive Thermopile Detector
AD8629 single op-amp analog gain circuit
Customer supplied PRT for heat sink
monitoring



### **Signal Conditioning Electronics**

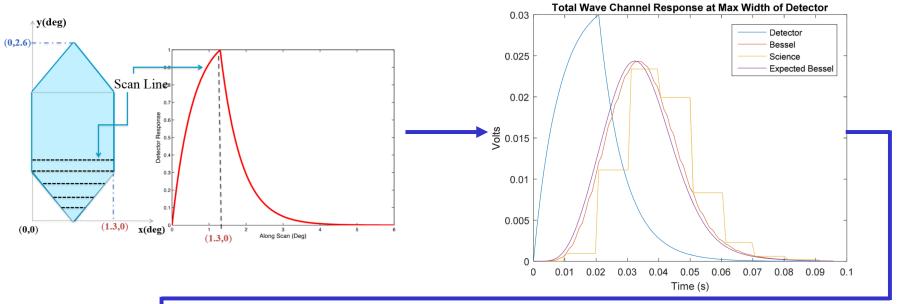


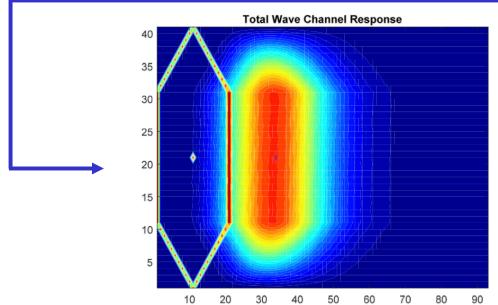




## **Signal Conditioning Electronics**





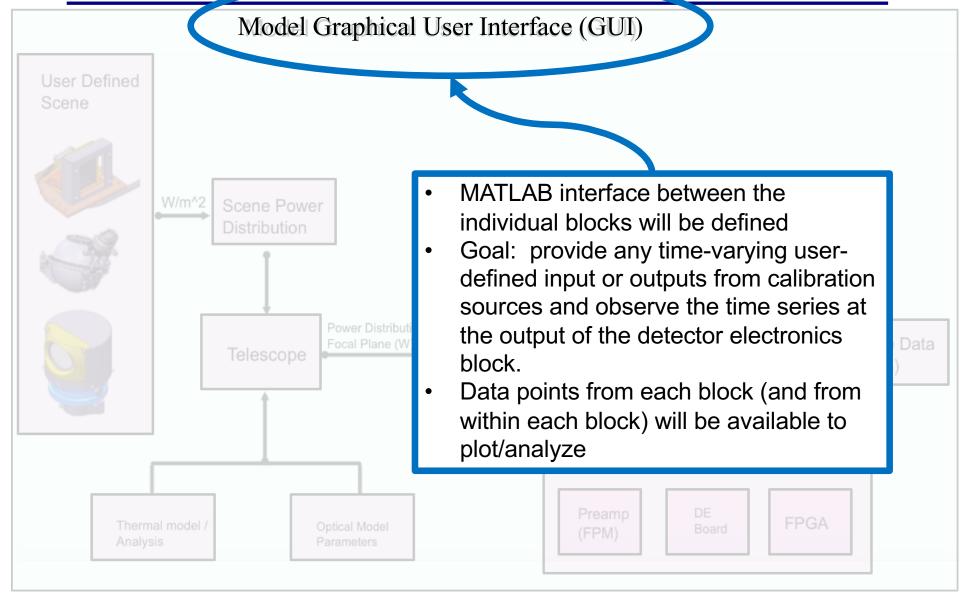


PSF Simulation



### **Graphical User Interface**







### **Current Status and Future Work**



- Currently in Build 1 phase: All subassemblies are being developed in their respective platforms.
  - ✓ Design changes are being incorporated as engineering drawings become available
- Electronics model is nearly complete to the current design specifications
- Scene generator between calibration targets and telescope currently being defined and developed
- On-going thermal analysis supports and validates contractor's derived requirements for individual subsystems (ICT, telescope)
- Short-term studies that can influence instrument design are also being carried out in parallel
  - ✓ Stray light studies
  - ✓ SW filter heating and re-emission
  - ✓ Temperature variations in telescope baffles due to material change
  - ✓ Uncertainties in radiance arriving at telescope aperture due to:
    - View angles for all three telescopes to the sources.
    - Uncertainties in knowledge of the system parameters- ICT temp, paint absorptivities, BRDFs, etc.

# Questions?

